博士論文 (要約)

Geochemical cycles of volatile elements in surface environment of the Earth

(地球表層における揮発性元素の物質循環)

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Abstract

Volcanic and hydrothermal activity discharge variable volatile species from the Earth's mantle to surface environment. Sulfur and carbon are major components of volcanic gases and they significantly affect atmospheric and oceanic chemistry. Halogens form variable chemical compounds in surface environments due to their high chemical activity. Therefore their fluxes are important to reveal Earth's evolution and to assess how much volcanic activity affects chemistry of surface environment. For sulfur and halogen fluxes, an approach using calibration against the reliable ³He flux was applied for elucidating their cycles. Especially for chlorine, isotope compositions were determined using the new experimental method with high resolution applicable for microinclusions. Helium and carbon behavior related to eruption was investigated based on gas chemistry.

In order to estimate sulfur flux at mid-ocean ridges (MOR), I focused on the volatile components preferentially released from magmas through film rupturing processes in ascending magmas. Such components are expected to be retained in vesicles of MORB glasses, and they are analyzed with the currently established crushing technique based on the conventional noble gas analyses. Sulfur fluxes were calibrated against ³He using datasets of MORB vesicles, submarine hydrothermal fluids, and volcanic gases. Carbon fluxes were also calibrated against ³He fluxes based on MORB and volcanic gas datasets. The obtained sulfur flux at MOR is less than their volcanic fluxes at arc volcanoes (ARC). Wedge mantle contributions to ARC fluxes were deconvolved, and the mantle fluxes combining MOR and ARC were calculated for sulfur and carbon. The estimated C/S flux ratio is close to the C/S ratio in the present surface inventory, suggesting their mantle origins. To balance the mass between surface environment and the mantle, fluxes of sulfur and carbon subducted into the mantle and not immediately returned to the surface were assessed.

Halogen cycles were also discussed. In order to estimate halogen fluxes at MOR, I focused on the volatile components preferentially released from magmas as sulfur, and analyzed MORB glasses. The complete halogens and helium datasets in MORB vesicles were obtained by combinations of some independent analyses, providing halogen fluxes at MOR by calibrating against the well documented ³He flux.

This part will be published in five years.

この部分については、5年以内に雑誌等で刊行予定のため非公開。

Geochemical cycles of variable volatile elements and approaches to investigate them are described in this dissertation.

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Chapter 1 General introduction

1. Volatile flux from the mantle and evolution of atmosphere and ocean

"How surface environment of the Earth was constructed" is one of the long-lasting theme of geoscience. Volcanic and hydrothermal activity discharges volatile elements such as sulfur, carbon, and halogens from the Earth's mantle to the surface, which has formed atmosphere and ocean. Therefore in order to constrain surface environmental evolution, it is important to assess volatile behavior between the interior and the surface of Earth and its degassing histories. In particular, estimations of volatile flux from the mantle to the surface enable us to calculate residence times of volatiles in surface environment as follows:

$$\mathbf{T}_{\mathbf{V}} = \mathbf{S}_{\mathbf{V}} / \mathbf{F}_{\mathbf{V}} \tag{1-1}$$

where the subscript "V" shows a target volatile element; T_V, S_V, and F_V indicate the residence time, the surface inventory at present, and the mantle flux of V, respectively. We may discuss Earth's degassing and atmospheric evolutionary histories by comparing calculated residence times of volatiles and the Earth's age of 4.54×10^9 years (Tera, 1981). When residence times are longer than the Earth's age, they imply catastrophic degassing histories of the volatile elements, which is consistent with the argon degassing history (Hamano & Ozima, 1978; Pujol et al., 2013; briefly depicted in Fig. 1-1). The argon degassing and evolutionary history has been well constrained based on argon isotopic compositions in the atmosphere and the mantle. One might calculate the duration of the early catastrophic degassing event of argon and the degassing fraction using the firstorder rate box model focusing on mantle degassing and argon isotopic composition in the atmosphere (Hamano & Ozima, 1978). Parameters used in the model were isotopic compositions and abundance of argon in the atmosphere and the mantle, potassium contents in the crust and the mantle, decay constants of ⁴⁰K which produces ⁴⁰Ar as a daughter isotope, masses of the crust and the mantle, and the degassing rate. They concluded that more than ~80% of 36 Ar in today's atmosphere was discharged by 4.0×10^9





The brown solid curve indicates minimum degassing fractions of ³⁶Ar contained in the Earth. The minimum degassing fractions for 0-0.9 Gy are values estimated by Hamano & Ozima (1978), while those for 0.9-4.54 Gy are estimated values under the assumption that 80% of ³⁶Ar remaining in the Earth has been degassed continuously during the past 3.64×10^9 years.

years ago. The growth history of the continental crust was recently assessed based on argon analyses of Archean hydrothermal quartz (3.5 Ga) and the atmospheric argon evolution model of Hamano & Ozima (1978), suggesting the catastrophic mantle degassing during the Earth's first 170 My and formation of the significant volume of the crust comparable to 80+/-10% of its present abundance between 3.8 Ga and 2.5 Ga (Pujol et al., 2013). The catastrophic mantle degassing event in the early Earth is widely accepted in the geoscience field.

Volatiles are injected from the mantle to the surface through submarine and land volcanisms. In the present Earth, the main mantle flux of volatile elements seems to be derived from submarine volcanisms related to mid-ocean ridges (MOR), which are the dominant divergent plate boundaries with the length of 1.5 times (Uyeda & Kanamori, 1979) longer than that of the total arc trench at convergent boundaries. Volatile flux at MOR has been determined mainly based on volatile contents in submarine glasses and the production rate of oceanic crust as follows:

$$F_{V} = F_{MORB} x [V]$$
(1-2)

In this equation, the target volatile is indicated as "V". F_V and F_{MORB} indicate flux of V and the production rate of MORB (MOR basalt, i.e., oceanic crust), respectively. [V] indicates V concentration in the carrier (i.e., MORB glass here). In this case, flux is calculated under the assumption that the all the volatile in the oceanic crust are finally released into the ocean. Therefore in the case of volatile elements which tends to stay in the oceanic crust and is not easily released into the ocean, more accurate MOR flux is estimated as follows:

$$F_{V} = F_{MORB} x [V] x R_{V}$$
(1-3)

This equation is derived from modification of eq. 1-2. F_V , F_{MORB} , and [V] are the same as those explained for eq. 1-2, and R_V is a ratio of the volatile "V" amount released from the mantle into the ocean relative to the total V amount in the oceanic crust. However quantifying R_V is not easy for variable elements because it may be succeeded by isotope exchange-based calculations, which requires high precision isotope analyses and cannot

be applied for elements with only one stable isotope such as fluorine. When a precise volatile flux from the mantle cannot be easily calculated multiplying the MORB flux and volatile concentrations therein, another possible way is the calculation based on a well determined flux of a specific volatile element as follows:

$$F_{V1} = F_{V2} x [V1]/[V2]$$
 (1-4)

V1 is the target volatile element, while V2 is a volatile element whose mantle flux is well determined somehow. Fv1 and Fv2 are mantle fluxes of V1 and V2, respectively. [V1] and [V2] indicate respectively V1 and V2 concentrations in the same reservoir (i.e., mantle derived materials at ridges in this equation). ³He can be conventionally applied as V2 because its MOR flux is well determined based on excess ³He in deep seawater and ³He/CO₂ ratios in fresh MORB (e.g., Craig et al., 1975; Jean-Baptiste, 1992; Saal et al., 2002; Cartigny et al., 2008; Bianchi et al., 2010). In particular, high ³He excess in deep seawater can be explained only by ³He injection from the mantle because it is primordial in origin and is not produced by radioactivity efficiently. Thus ³He anomaly in deep seawater well reflects the 3 He flux at ridges. The recent study reported the MOR 3 He flux of (527+/-102) mol/y comparing World Ocean Circulation Experiment datasets of helium isotopes in deep seawater with depths greater than 1000m and calculated ³He and radiocarbon compositions in deep seawater based on a suite of ocean general circulation models (Bianchi et al., 2010), which is a half value of $(1070+/-270)^{3}$ He mol/y (Craig et al., 1975) conventionally accepted. In addition, the global ³He flux at volcanic arcs (ARC) is calculated to be (105+/-20) mol/y as the 20% value of MOR ³He flux (Torgerson, 1989). We therefore obtain the global ³He cycle (Fig. 1-2). This obvious geochemical cycle and inert chemistry of noble gases enable ³He to be one of the most useful geochemical tracers, and its mantle flux has been conventionally used for calculations of other volatile fluxes as follows:

$$F_V = F_{helium-3} x [V]/[^3He]$$
 (1-5)

This equation is a modified version of eq. 1-4. F_V and $F_{helium-3}$ are fluxes of volatile "V" and ³He, respectively. This equation can be applied for flux calculations when V/³He



Figure 1-2. Global ³He fluxes among the Earth's interior and surface.

The mid-ocean ridge flux is the value reported by Bianchi et al. (2010). The flux from arc volcanoes was calculated multiplying MOR flux by 20% which is the arc flux ratio relative to MOR flux (Torgersen, 1989). Subscripts indicate errors of the fluxes. The hotspot flux is the value reported by Torgersen (1989). The subducting ³He flux is the value reported by Hilton et al. (2002) under the assumption that there is only terrestrial ³He in marine sediment. The residence time of ³He is the half of that of ⁴He (~1 My; Ozima & Podosek, 2002). Atmospheric ³He continuously escapes into the space due to its small atomic weight. Helium-3 fluxes could be easily determined because its behavior is basically unidirectional from the mantle to surface.

ratios are not affected by fractionation through degassing procedures. [V] and [³He] are concentrations of V and ³He in the same reservoir such as MORB glasses. In particular, carbon flux has been well constrained by calibrating against ³He flux (e.g., Marty & Jambon, 1987; Marty & Tolstikhin, 1998). Mantle fluxes and geochemical cycles of volatile elements will be estimated based on the calculation methods explained in the current section.

2. Research objectives and strategies

In this research, for assessing how Earth's degassing has been producing atmosphere and ocean, I estimate global mantle fluxes of sulfur, carbon, and halogens which form various chemical compounds and significantly affect surface environment on the Earth. Based on the estimated mantle fluxes, their geochemical cycles and degassing histories are also discussed. This study constrains volatile geodynamics and degassing histories based on the well accepted ³He fluxes in the surface.

In *Chapter 2*, I discuss sulfur and carbon geochemical cycles and degassing histories. The currently established crushing technique enables direct comparison of sulfur and helium retained in MORB vesicles, which may be a component preferentially released at MOR. Applying data of MORB and hydrothermal fluids around the world, sulfur and carbon fluxes at MOR are calculated in this chapter. In addition, the wedge mantle contribution to volcanic gases is deconvolved in order to calculate the mantle fluxes at ARC. Comparing MOR and ARC fluxes, geochemical cycles and degassing histories of sulfur and carbon are constrained.

In *Chapter 3*, I discuss halogen geochemical cycles and degassing histories. The crushing method used in *Chapter 2* is applied for comparing fluorine, chlorine, and helium in MORB vesicles directly. In addition, the neutron irradiation method converting halogens into proxy noble gas isotopes is also applied for calculating halogen concentrations and inter-halogen ratios in MORB glasses. Based on these datasets, mantle fluxes of halogens are estimated in this chapter. Comparing MOR and ARC fluxes, halogen geochemical cycles and degassing histories are constrained.

Chapter 2 Geochemical cycles of sulfur and carbon

1. Introduction

Volcanic and hydrothermal activity discharges variable volatile species from the Earth's mantle to surface. Sulfur and carbon, which are both major components of volcanic gases, are injected into the surface in forms of SO₂, H₂S, CO₂, and CH₄, and significantly affect atmospheric and oceanic chemistry and their evolution. In order to reveal Earth's evolution and to assess how much volcanic activity affects chemistry of surface environment, previous studies were trying to estimate sulfur and carbon fluxes between the Earth's interior and the surface. However, only a small number of studies have discussed global mass balances of sulfur and carbon (Hilton et al., 2002; Wallace, 2005; see reported sulfur fluxes in Table 2-1) while other studies have focused only on local ridges and arcs. It should be noted that elucidating individual volcanic and hydrothermal activity is very important for assessing their effects on human society (hazard mitigation), and is also essential for estimating global volatile cycles by comparing themselves as representatives of volcanic activity around the world. I will discuss sulfur and carbon cycles among Earth's surface and interior reservoirs based on chemistry of ridges and arcs worldwide.

In surface environment, sulfur and carbon are stored as gases (SO₂, H₂S, CO₂, CH₄ etc.) in the atmosphere, dissolved matters (SO₄²⁻, HCO₃⁻ CO₃²⁻ etc.) in seawater, and are also retained as solid materials (FeS₂, CaCO₃ etc.) in sedimentary reservoirs. They finally move to the mantle owing to subduction of oceanic plates at trenches, which makes a part of them recycled back into the mantle (e.g., Hilton et al., 2002), while other parts are returned to the atmosphere and ocean again. It is important to estimate sulfur and carbon fluxes between the mantle and the surface for revealing histories of Earth's degassing and atmospheric evolution. In particular, when the mantle flux of sulfur is obtained, applying eq. 1-1, I can estimate the residence time of sulfur. Thus I can discuss the sulfur degassing history by comparing the calculated residence time and the Earth's age of 4.54×10^9 years (Tera, 1981).

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S flux [Gmol/y]	Residence time [Gy]	Flux estimation method	Ref.
2670	0.2	Loss of total sulfur in MORB	Hilton et al. (2002)
1640	0.3	Loss of total sulfur in MORB	Hansen & Wallmann (2003)
144	3.7	Alteration of ~10% of MORB	Sakai & Matsuhisa (1996); Tajika (1998)
78	6.8	Based on altered oceanic crust compositions	Alt (1995)

Table 2-1. Reported sulfur global fluxes at ridges and residence times.

NOTE: Sulfur surface inventory of 5.3×10^{20} mol (Hilton et al., 2002) was applied for residence time calculations.

The global flux of sulfur from sub-aerial arc volcanoes is well constrained (Hilton et al., 2002; Bluth et al., 1993) based on measurements of SO₂ gas from volcanoes using correlation spectrometry abbreviated as COSPEC (Stoiber & Jepsen, 1973), UV spectroscopy, and satellite remote sensing (Bluth et al., 1993). Carbon flux from volcanic arcs can be also obtained multiplying global SO₂ flux and a representative CO₂/SO₂ ratio in volcanic gases, or following another estimation method discussed later. However, the main mantle flux of volatile elements seems to be derived from submarine volcanisms related to MOR as discussed in Chapter 1. Therefore, we need to discuss MOR flux together with sub-aerial flux to evaluate a mass balance of sulfur and carbon on the Earth's surface precisely. However MOR volatile flux cannot be determined by direct observations in ways similar as those for arc volcanoes such as COSPEC and satellite remote sensing because the measurement of ocean-bottom hydrothermal systems is technically very difficult to succeed. It seems that MOR fluxes can be calculated based on volatile concentrations in MORB and the MORB production rate (eq. 1-2), but such bulk concentration-based flux is often inconsistent with the catastrophic degassing history of Earth expected from atmospheric argon evolution (Hamano & Ozima, 1978; Pujol et al., 2013; Fig. 1-1). In the case of sulfur, based on its concentration in the oceanic crust, fluxes of 2.67×10^{12} mol/y and 1.64×10^{12} mol/y at MOR were evaluated by Hilton et al. (2002) and Hansen & Wallmann (2003), respectively. Applying these fluxes and the sulfur surface inventory of 5.33×10^{20} mol (Hilton et al., 2002) to eq. 1-1, residence times of sulfur in the surface reservoir are calculated to be 2.0×10^8 years and 3.3×10^8 years, respectively. These residence times are significantly shorter than the Earth's age of 4.54×10^9 years, which seems to imply that there were no need to have a significant degassing event in the early Earth. It seems, however, that if only a part of total sulfur in the oceanic crust is finally released into the ocean by alteration of the crust, the net sulfur flux from the mantle to the ocean is lower than the flux calculated using sulfur contents in the oceanic crust and the MORB flux. This is why we have to calculate MOR flux of volatiles including sulfur more precisely in another way.

One of ways to estimate volatile flux from the mantle very precisely is evaluating isotopic exchange between oceanic crust and seawater. A sulfur flux of 7.8×10^{10} mol/y from the volcanic section in the oceanic crust to the ocean was calculated from seawater–basalt sulfur exchange during hydrothermal alteration based on sulfur isotope

measurements in Ocean Drilling Program Hole 504B samples (Alt, 1995). This flux is markedly lower than the estimate of 2.67×10^{12} mol/y and 1.64×10^{12} mol/y discussed above (Hilton et al., 2002; Hansen & Wallmann, 2003). This isotope exchange-based flux is very close to the flux of 1.44×10^{11} mol/y estimated under the assumption that only 10% of sulfur retained in the oceanic crust are finally released into the ocean (Sakai & Matsuhisa, 1996). Though the 10%-release of sulfur from the oceanic crust (i.e., $R_V = 0.1$ in eq. 1-3) is just an assumption and is not a value derived from a rigorous estimate, this flux may be much more reliable as a net sulfur flux relative to those calculated under the assumption that all the sulfur in the oceanic crust are released into the ocean. However, quantifying the reliable R_V is very difficult as discussed in *Chapter 1*, and more data for more sampling sites are needed to investigate the global sulfur flux precisely.

In this study, I will estimate sulfur flux in a different way. A precise sulfur flux from the mantle cannot be easily calculated multiplying the MORB flux and sulfur concentrations therein, thus I will calculate it by calibrating against the well determined ³He flux at MOR (Bianchi et al., 2010; Fig. 1-2; eq. 1-5). Carbon flux are also constrained in the same way using ³He flux (e.g., Marty & Jambon, 1987; Marty & Tolstikhin, 1998). Therefore mantle fluxes and geochemical cycles of sulfur and carbon will be discussed based on ³He flux in this chapter.

2. Carbon cycle

Carbon fluxes from the mantle to the atmosphere and the ocean have been already documented, but the conventional estimate (Marty & Talstikhin, 1998) depends on the old ³He flux from the mantle reported in 1970's (Craig et al., 1975). The MOR ³He flux has been recently modified (Bianchi et al., 2010), thus one might recalculate global carbon fluxes based on the newly estimated ³He flux from the mantle. The method to estimate global carbon cycles will be also useful for investigating sulfur cycles.

2.1. Carbon flux at mid-ocean ridges

Mantle carbon flux of $2x10^{12}$ mol/y was derived from the MOR ³He flux and C/³He ratio in MOR basalt glass (Marty & Jambon, 1987), but the most accepted carbon flux at MOR was estimated by Marty & Tolstikhin (1998). They calculated carbon flux based on C/³He ratios in submarine glasses and hydrothermal fluids at MOR. They treated their C/³He ratios as a single population and obtained the average C/³He ratio of $(2.2+/-0.7)x10^9$. This ratio was multiplied by the MOR ³He flux of (1000+/-250) mol/y derived from a single oceanic model (Craig et al., 1975; Welhan & Craig, 1983), providing MOR carbon flux of $(2.2+/-0.9)x10^{12}$ mol/y (Marty & Talstikhin, 1998). However, the ³He flux has been recalculated to be (527+/-102) mol/y (Bianchi et al., 2010) based on the data and models more sophisticated than those used in Craig et al. (1975), which requires reestimation of carbon flux at ridges. Now, one might calculate carbon flux to be $(1.2+/-0.4)x10^{12}$ mol/y (Fig. 2-1) multiplying the representative upper mantle C/³He ratio of $(2.2+/-0.7)x10^9$ and the MOR ³He flux of (527+/-102) mol/y (Bianchi et al., 2010).

2.2. Carbon flux at arc volcanoes

The carbon and helium behaviors in individual volcano are well studied (see *Appendix C* for an example at Ontake volcano). Following the same way as the MOR flux calculation, the global carbon flux at arc volcanoes can be estimated multiplying $C/{}^{3}$ He ratios in volcanic gases and the determined 3 He flux of (105+/-20) mol/y at ARC (Fig. 1-2). There are many subaerial volcanoes in the circum-Pacific region due to tectonic settings of convergent plate boundary (Fig. 2-2; see also Table 2-2). Helium and carbon isotopic compositions in high temperature (>200C) volcanic gases and steam wells are



Figure 2-1. Global carbon fluxes among the Earth's interior and surface environment.

Numbers are carbon fluxes among reservoirs with errors as subscripts. They are described in the unit of Gmol/y. See text for calculations of fluxes at mid-ocean ridges (1200 Gmol/y) and arc volcanoes (2100 Gmol/y), and the mantle contribution to the ARC flux (230 Gmol/y). Other fluxes related to recycle and return were calculated under the assumption of the steady-state surface environment of carbon.



Figure 2-2. Sites of MORB, submarine hydrothermal fluids, and arc volcanoes considered in this study.

Stars, circles, and triangles indicate sites of MORB glasses, submarine high temperature (>200C) hydrothermal fluids, and high temperature (>200C) volcanic gases, respectively. Names of each site and sample are described in Table 2-2. This figure was prepared using the Ocean Data View software (Schlitzer, 2013).

Number in Fig. 1-4	Sample/site name	Lat.	Lon.	Depth [m]
(MORB)				
M1	CL DR01	12.75oN	103.93oW	2610
M2	ND 18-1	17.42oS	113.20oW	2578
M3	RY380-R03b	17.44oS	113.22oW	2580
M4	CH31 DR12	36.870N	33.25oW	2895
M5	CH98 DR12	30.17oN	41.92oW	4150
M6	418R002	15.36oN	46.63oW	4944
M7	KH10-6 DR15-glass	24.48oS	69.92oE	2242
M8	KH10-6 DR16-glass	24.53oS	69.93oE	2268
(Submarine hydrothermal	fluids >200°C)			
H1	Juan de Fuca Ridge (47N-48N)	48.0N	129.1W	
H2	East Pacific Rise (21N)	20.8N	109.1W	
H3	East Pacific Rise (11N-13N)	11.0-12.8N	103.5-104.0W	
H4	East Pacific Rise (17S-19S)	17-19S	113.5W	
H5	Mid-Atlantic Ridge (36N-38N)	36.2-37.8N	31-33.9W	
H6	Mid-Atlantic Ridge (23N-26N)	23.4-26.1N	45W	
(Arc volcanoes)				
A1	Klyuchevskoy	56.06N	160.64E	
A2	Koryak	53.32N	158.69E	
A3	Avacha	53.26N	158.83E	
A4	Mutnovsky	52.45N	158.2E	
A5	Chirpoy	46.5N	151E	
A6	Kudryavy	46N	148E	
A7	Usu	42.53N	140.83E	
A8	Kuju	33.1N	131.3E	
A9	Unzen	32.75N	130.3E	
A10	Satuma-Iwojima	30.8N	130.3E	
A11	Merapi	7.52S	110.43E	
A12	Lewoto	8.27S	123.5E	
A13	Ngawha	35.4S	173.85E	
A14	White Island	37.52S	177.18E	
A15	Ohaaki	38.51S	176.29E	
A16	Mokai	38.52S	175.91E	
A17	Ngauruhoe	39.15S	175.63E	
A18	La Primavera	20.62N	103.52W	
A19	Colima	19.50N	103.62W	
A20	Pacaya	14.37N	90.6W	
A21	Cerro Negro	12.5N	86.7W	
A22	Momotombo	12.41N	86.52E	
A23	Galeras	1.22N	77.37W	
A24	Cumbal	0.95N	77.89W	

Table 2-2. Sample or site names considered for flux calculations in Chapter 1.

NOTE: Numbers correspondent to those in Fig. 2-2. Names of MORB are sample names, while others are site names.

is in the circum-Pacific region.	He Reference		+10 Taran et al. (1991); Taran (2009)	+10 Taran et al. (1997); Taran (2009)	+10 Taran et al. (1997); Taran (2009)	+10 Taran et al. (1992); Taran (2009)	+10 Taran et al. (1995); Taran (2009)	+09 Fischer et al. (1998); Taran (2009)	+10 Marty et al. (1989)	+10 Nagao et al. (1981); Saito et al. (2002)	+09 Sano & Williams (1996)	+09 Marty et al. (1989); Taran (2009)	+10 Javoy et al. (1986); Taran (2009)	+10 Varekamp et al. (1992)	+10 Giggenbach et al. (1993);Lyon and Hulston (1983)	+10 Marty & Giggenbach (1990)	+10 Giggenbach et al. (1993)	+10 Giggenbach et al. (1993)	+09 Giggenbach et al. (1993); Giggenbach (1996); Fischer (2008)	+09 Taran et al. (2002)	+09 Sano & Williams (1996); Taran (2009)	+10 Sano & Williams (1996)	+10 Sano & Williams (1996)	+10 Javoy et al. (1986); Taran (2009)	+10 Sano & Williams (1996)	+10 Sano & Williams (1996)
nal fluid	C/ ³		1.1E	3.6E-	2.3E-	1.3E	5.9E-	9.0E	1.0E	1.4E [.]	8.1E [.]	6.3E [.]	1.4E [.]	1.3E	2.8E [.]	3.6E-	6.9E [.]	1.3E	7.9E	6.2E	6.1E	2.0E-	3.3E [.]	1.6E [.]	1.5E [.]	1.2E [.]
) geotheri	$\delta^{13}C$	(%)	-11.6	-11.8	-5.7	-9.9		-7.2	-4.4	-7.9	-6.1	-5.5	-3.9	-3.9	-8.7	-2.0			-10.4	-3.9	-6.3	-6.9	-2.5	-2.8	-7.6	-4.9
ture (>200C	³ He/ ⁴ He	(Ratm)	6.8	5.3	7.1	8.2	6.3	6.8	5.4	7.6	7.3	7.9	7.4	3.6	6.0	6.1	5.0	6.8	6.7	6.2	6.6	4.2	6.8	7.1	8.4	6.8
gh tempera	Temp.	(°C)	250	220	473	543	344	920	750	351	803	885	803	490	269	495	280	300	516	260	814	965	350	747	222	257
compositions in hi	Location		Kamchatka	Kamchatka	Kamchatka	Kamchatka	Kuril	Kuril	Japan	Japan	Japan	Japan	Indonesia	Indonesia	New Zealand	New Zealand	New Zealand	New Zealand	New Zealand	Mexico	Mexico	Guatemala	Nicaragua	Nicaragua	Colombia	Colombia
isotopic o	Lon.		160.64E	158.69E	158.83E	158.2E	151E	148E	140.83E	131.3E	130.3E	130.3E	110.43E	123.5E	173.85E	177.18E	176.29E	175.91E	175.63E	103.52W	103.62W	90.6W	86.7W	86.52E	77.37W	77.89W
and carbor	Lat.		56.06N	53.32N	53.26N	52.45N	46.5N	46N	42.53N	33.1N	32.75N	30.8N	7.52S	8.27S	35.4S	37.52S	38.51S	38.52S	39.15S	20.62N	19.50N	14.37N	12.5N	12.41N	1.22N	0.95N
Table 2-3. Helium a	Name		Klyuchevskoy	Koryak	Avacha	Mutnovsky	Chirpoy	Kudryavy	Usu	Kuju	Unzen	Satuma-Iwojima	Merapi	Lewoto	Ngawha	White Island	Ohaaki	Mokai	Ngauruhoe	La Primavera	Colima	Pacaya	Cerro Negro	Momotombo	Galeras	Cumbal

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Figure 2-3. Carbon and helium compositions in volcanic gases and their end-members. Volcanic gas data described in Table 2-3 and 2-4 are shown as red circles. Three square zones are end-members of volcanic gases: The upper mantle component; organic sediment; and limestone with slab carbonate abbreviated as "limestone". Mixing lines of three end-members are depicted as dashed curves. Dashed lines are derived by the medium upper mantle composition and medium δ^{13} C values of organic sediment and limestone. This figure shows that all the data of volcanic gases are explained by mixing of the three end-members as suggested by Sano & Marty (1995).

compiled in Table 2-3. An average value of C/³He ratios in 24 subaerial volcanoes in Table 2-3 is calculated to be $(2.0+/-0.3)x10^{10}$, providing the ARC carbon flux of $(2.1+/-0.5)x10^{12}$ mol/y by multiplication of the ³He flux of (105+/-20) mol/y and the C/³He ratio of $(2.0+/-0.3)x10^{10}$ at ARC. This ARC flux is consistent with the recent estimate using volcanic gas observations worldwide (Fischer, 2008).

2.3. Deconvolution of carbon flux at ARC

In order to estimate the carbon flux from the mantle to the surface through arc volcanisms, I deconvolved magmatic carbon into the mantle component and other endmembers. Volatiles in volcanic gases are derived from magmas which have mainly two sources: components derived from the mantle and subducting materials. Deconvolution of the source contributions to volcanic gases is important to estimate net volatile fluxes from the mantle. In particular, carbon and helium mixing at subduction zone has been well studied. Sano & Marty (1995) compared carbon and helium isotopic compositions in surface reservoirs and the mantle, and concluded that those isotopic compositions in volcanic gases are produced by three-component mixing: the upper mantle component represented by MORB, organic sediment, and marine limestone with slab carbonate (abbreviated as "limestone" hereafter) (see also Fig. 2-3). Following Sano & Marty (1995), δ^{13} C values of -6.5+/-2.5 ‰ in the upper mantle, -30+/-10 ‰ in organic sediment, and 0+/-2 ‰ in limestone were used as end-member compositions. The C/³He ratio in the upper mantle is (2.2+/-0.7)x10⁹ (Marty & Talstikhin, 1998). In addition, Sano & Marty applied the $C/{}^{3}$ He ratio of $10^{(13+/-1)}$ for subducting material end-members (organic sediment and limestone) based on the compositions of crustal CO₂-rich continental gases (O'Nions & Oxburgh, 1988), thus I also used the same $C/{}^{3}He$ ratio for an end-member value. Volcanic gas and end-member compositions are described in Table 2-4. Based on those values, respective contribution percentages of the three components in the ARC samples were calculated solving the following equations:

$$\delta^{13}C_{\rm V} = \delta^{13}C_{\rm M}M + \delta^{13}C_{\rm S}S + \delta^{13}C_{\rm L}L \tag{2-1}$$

$$1/(C^{3}He)_{V} = M/(C^{3}He)_{M} + S/(C^{3}He)_{S} + L/(C^{3}He)_{L}$$
 (2-2)

M + S + L = 1 (2-3)

Name	$\delta^{13}C$	C/ ³ He	Carbon source contribution (%)					
	(‰)		Upper mantle	Organic sediment	Limestone			
(High temperature volcanic gase	es and ste	eam wells)						
Klyuchevskoy	-11.6	1.1E+10	20.0	34.3	45.7			
Koryak	-11.8	3.6E+10	6.1	38.0	55.9			
Avacha	-5.7	2.3E+10	9.5	16.8	73.7			
Mutnovsky	-9.9	1.3E+10	16.9	29.3	53.8			
Chirpoy		5.9E+10	3.7					
Kudryavy	-7.2	9.0E+09	24.4	18.7	56.9			
Usu	-4.4	1.0E+10	22.0	9.9	68.1			
Kuju	-7.9	1.4E+10	16.0	22.9	61.1			
Unzen	-6.1	8.1E+09	27.3	14.2	58.4			
Satuma-Iwojima	-5.5	6.3E+09	34.9	10.8	54.3			
Merapi	-3.9	1.4E+10	15.7	9.6	74.7			
Lewoto	-3.9	1.3E+10	17.5	9.3	73.3			
Ngawha	-8.7	2.8E+10	7.8	27.2	65.1			
White Island	-2.0	3.6E+10	6.1	5.5	88.4			
Ohaaki		6.9E+10	3.2					
Mokai		1.3E+10	16.6					
Ngauruhoe	-10.4	7.9E+09	27.7	28.7	43.6			
La Primavera	-3.9	6.2E+09	6.6	7.0	86.4			
Colima	-6.3	6.1E+09	13.7	6.2	80.1			
Pacaya	-6.9	2.0E+10	11.2	20.7	68.1			
Cerro Negro	-2.5	3.3E+10	14.4	22.2	63.3			
Momotombo	-2.8	1.6E+10	18.8	12.3	68.9			
Galeras	-7.6	1.5E+10	36.1	13.2	50.8			
Cumbal	-4.9	1.2E+10	35.5	5.3	59.2			
Average values	-6.4	2.0E+10	11.0	18.9	70.1			
(End-members)			_	_				
Upper mantle	-6.5	2.2E+09	100.0					
Organic sediment	-30.0	1.0E+13		100.0				
Limestone with slab carbonate	0.0	1.0E+13			100.0			

Table 2-4. Sample and end-member compositions with carbon contributions.

Data of subaerial volcanoes and three end-members are described here. See text for the calculations of end-member contributions.

where subscripts M, S, and L indicate contributions of the upper mantle component, organic sediment, and limestone, respectively. Subscripts V, M, S, and L of δ^{13} C values and $C/^{3}$ He ratios correspond to compositions of high temperature volcanic gases, the upper mantle, organic sediment, and limestone. Calculated contributions are described in Table 2-4. The contribution of the upper mantle carbon was 3.2%-36.1% with an average of (11.0+6.5/-4.7)%. The uncertainty of the upper mantle contribution was calculated based on the highest and the lowest upper mantle contributions calculated based on extremely high or low values' sample and end-member compositions as follows: the upper mantle contribution of 17.5% derived from the case applying C/³He ratios of 1.7×10^{10} , 2.9×10^9 , 1.0×10^{14} , and 1.0×10^{14} for ARC samples, the upper mantle, organic sediment, and limestone; and the upper mantle contribution of 6.3% derived from the case applying C/³He ratios of 2.3×10^{10} , 1.5×10^{9} , 1.0×10^{12} , and 1.0×10^{12} for ARC samples, the upper mantle, organic sediment, and limestone. This upper mantle contribution indicates that ~90% of carbon released at arc volcanoes is attributable to subducted materials, that is, carbonate and organic carbon. As discussed above, the global carbon flux at ARC is $(2.1+/-0.5)x10^{12}$ mol/y. The end-member deconvolution implies that 11% of carbon in high temperature volcanic gases is derived from the upper mantle component, providing carbon flux of $(2.3+1.5/-1.2)\times 10^{11}$ mol/y from the mantle to atmosphere via arc volcanoes.

2.4. Contribution to global flux of carbon from present hot spot magmatism

C/³He ratios in hotspot areas have been well observed. Marty and Jambon (1987) calculated an average C/³He ratio of 5.3×10^9 in Loihi samples using already reported data (Kurz et al., 1983; Rison & Craig, 1983; Exley et al., 1986). They also determined the C/³He ratio in Kilauea gases as 3.4×10^9 based on observational datasets (Naughton et al., 1973; Craig & Lupton, 1976). In addition, C/³He ratios of 2.1×10^9 and 2.7×10^9 are reported for Yellowstone (Craig et al., 1978), and 2.0×10^9 and 7.9×10^9 are reported for Iceland gases (Sano et al., 1985; Mamyrin & Tolstikhin, 1984). These values are slightly higher than the MORB value. However expected hotspot C/³He ratios are lower than MORB because their deep sources are less degassed relative to the MORB mantle, so the observed data should be treated with caution (Marty & Jambon, 1987). In addition, ³He flux at MOR is 200 times higher than that at hotspots, therefore the hotspot contribution to the global carbon flux is negligibly small.

2.5. Global carbon cycle

A summary of global carbon flux is depicted in Fig. 2-1. The total volcanic flux of carbon is $(3.3+/-0.6)\times10^{12}$ mol/y, which is two orders of magnitude smaller than anthropogenic emission of $(6.9+/-0.6)\times10^{14}$ mol/y by fossil fuel combustion and cement production (Ciais et al., 2013). The MOR flux combined with the wedge mantle flux is $(1.4+/-0.4)\times10^{12}$ mol/y. This value, if accumulated for 4.54×10^{9} years, results in $(6.5+/-1.9)\times10^{21}$ mol of carbon, which is comparable to the surface inventory of 7.0×10^{21} mol (Hilton et al., 2002). If steady-state recycling of carbon is applied, then the total subduction flux becomes $(3.3+/-0.6)\times10^{12}$ mol/y. This estimate is consistent with the carbon subduction flux of 3.46×10^{12} mol/y (Hilton et al., 2002).

3. Sulfur cycle

3.1. How to assess upper mantle-derived compositions

One might estimate sulfur fluxes from the mantle to the surface based on ³He fluxes in the same way as carbon discussed in the previous section. However, except for a very rough estimate (Kagoshima et al., 2012), no report in the literature has described an estimation of mantle sulfur flux at MOR as calibrated against the mantle ³He flux because representative $S/^{3}$ He ratios released from the upper mantle into the ocean have never been reported.

In the case of the carbon and helium coupling, the representative material for the upper mantle is MORB glasses. They have C/³He ratios close to 2.0×10^9 both in CO₂ vesicles and glass matrix under the equilibria of helium in those two phases (Marty & Jambon, 1987), implying that degassing and volatile release from MORB does not cause significant fractionation of $C/{}^{3}$ He ratios therein. Therefore, $C/{}^{3}$ He ratios released from the upper mantle into the ocean can be represented by compositions of MORB glasses. It should be noted that the representative $C/{}^{3}$ He ratio in the upper mantle had been modified to be 2.2×10^9 (Marty & Talstikhin, 1998) as discussed in the previous section. On the other hand, it is much more complicated and difficult to determine S^{3} He ratios which represents the released components from the upper mantle into the ocean than representative $C/{}^{3}$ He ratios because $S/{}^{3}$ He ratios in MORB vesicles and glass matrix are much discrepant due to the difference between partition coefficients of sulfur and ³He between CO₂-H₂O vapor (vesicles) and basaltic melt (glass matrix). In MORB glasses, sulfur is mostly stored in glass matrix, while helium mostly stays in vesicles. This is why S^{3} He ratios in vesicles or matrix of MORB glasses are not representatives of components transferred from the upper mantle into the ocean, while they apparently reflect the upper mantle composition itself.

Namiki and Kagoshima (2014) suggested one dominant degassing process from magmas to surface reservoirs by decompression experiments of syrup including bubbles as an analogue of ascending magmas including voids in a volcanic conduit. They found appearance of large voids in the syrup during decompression. An expansion of one void deep in the syrup leads to another void expansion, which results in the propagation of voids upward to the syrup surface and significant outgassing. This upward propagation and degassing pathways are generated by bubble film-rupturing, causing efficient degassing from syrup to the surface (Fig. 2-4). Here, syrup and involved voids can be also treated as analogues of basaltic melts and CO_2-H_2O bubbles in them at MOR which finally become matrix and vesicles of MORB glasses after chilling. Such process generating chilled margin indicates that vesicle components are much more dominant for degassing from magmas and fluxes from the upper mantle relative to glass matrix components which are not easily dissolved into the ocean like fluids in vesicles. Thus we will take vesicle components as a representative value injected from the upper mantle to the ocean, and another representative value should be derived from compositions of submarine hydrothermal fluids which are generated components through seawater-basalt interactions and degassing processes from the upper mantle.

In order to estimate representative components released from the upper mantle, I conducted two approaches below:

1) Measurement of sulfur and helium compositions in vesicles of MORB glasses.

2) Determination of representative sulfur and helium compositions in high temperature (>200C) submarine hydrothermal fluids compiling data already reported.

Representative S/³He ratios derived from the upper mantle is estimated comparing values obtained by approaches 1) and 2). For 1), MORB glasses were crushed *in vacuo* with alkaline solutions, and analyzed by appropriate conventional methods. As carbon, representative upper mantle-derived S/³He ratios and δ^{34} S values enable us to quantify sulfur and helium behavior at ARC systems and the global mass balance of sulfur.





Figure 2-4. Expected bubble evolutions during magma ascending in volcanic conduits.

This figure was illustrated based on Figure 2 and Figure 4 of Namiki & Kagoshima (2014). Red squares and white circles indicate the same cell in a particular magma and voids in volcanic conduits, respectively. The magma is ascending and decompression is proceeding righter and righter. During magma ascent, voids are getting larger and rising in the cell, which results in the conflict among voids. Once conflicting bubble films rupture, the voids connect each other. These bubble rising and evolution associated with film-rupturing propagate from the deep magma to the surface as described in dotted vectors. Finally, evolved voids in the surface cause efficient outgassing from the magma to the atmosphere or the ocean as described in a red vector on the right. This efficient outgassing process may dominate volatile fluxes at MOR.

3.2. Methods

3.2.1. Glass vesicle

It is difficult to measure the abundance of sulfur species such as H₂S and SO₂ in vesicles of MORB glass together with ³He because such sulfur species are highly reactive and so sticky that they easily adhere to the inner wall of a vacuum crushing vessel and purification lines, which makes impossible to collect and measure them in a high yield enough to ignore decrease in concentrations and isotope fractionations, especially in static gas-MS systems. This is why I developed a gas-extraction method called as "Frozen Crushing Method" (FCM; see experimental details in *Appendix A*; Kagoshima et al., 2012). Sulfur and ³He concentrations in vesicles of MORB glasses were determined using this method. Subsequently, collected sulfate ion in the alkaline solution was converted into BaSO₄ precipitations by adding BaCl₂ solutions, and δ^{34} S values derived from vesicles were obtained with an elemental analyzer (vario PYRO cube; Elementar Analysensysteme, GmbH) coupled to an isotope-ratio mass spectrometer (Delta XP; Thermo Fisher Scientific Inc.) via an interface (ConFlo IV; Thermo Fisher Scientific Inc.) at the University of Ottawa.

3.2.2. Glass matrix

For some MORB glasses, finely crushed powders (<212µm) were combusted at the temperature of 1800C *in vacuo* with a tantalum oven, and volatiles including noble gases were extracted into the vacuum line. Thus the abundance of helium and ³He/⁴He ratios in powders were determined with a noble gas mass spectrometer (VG5400; Waters Corp.) at the Atmosphere and Ocean Research Institute (AORI), the University of Tokyo, in the same way as helium extracted from vesicles. Sulfur contents in glass matrix were measured with a secondary ion mass spectrometer (NanoSIMS; Cameca SAS, Gennevilliers, France) at AORI, whereas δ^{34} S values of some samples were obtained using an elemental analyzer isotope-ratio mass spectrometer system (Isoprime-EA; Isoprime Ltd.) at the University of Tsukuba (Maruoka et al., 2003).

3.3. Results

3.3.1. Vesicles

I analyzed MORB glass samples collected at eight sites on East Pacific Rise, Mid-Atlantic Ridge, and Central Indian Ridge (Fig. 2-2). The ³He/⁴He ratios and ³He contents of vesicles in MORB glasses range from 7.96 R_a to 9.57 R_a (where R_a is the atmospheric ratio of 1.382x10⁻⁶ (Sano et al., 2013)), and from 5.8x10⁻¹⁶ mol/g to 5.1x10⁻¹⁵ mol/g (Table 2-5), respectively. The values agree well with data presented in an earlier report (Graham et al., 2002). The samples show similar sulfur contents ranging from 3.3×10^{-8} mol/g to 1.7×10^{-7} mol/g in vesicles. The S/³He molar ratios in vesicles range from 1.8×10^{7} to 1.3×10^8 with an average of $(4.6 + / -1.3) \times 10^7$ (1 σ), which is a little more variable relative to sulfur contents, but similar in the order. This $S/^{3}He$ ratio will be used for assessing representative components released from the upper mantle. Evaluating δ^{34} S values of vesicle sulfur was impossible due to their low abundance less than the blank contribution from the filtering system inevitable for converting sulfur into BaSO₄ precipitations for gas-MS analyses. Thus I assume that the initial δ^{34} S value of 0% derived from the upper mantle (Marini et al., 2011) is also retained in vesicles. This value is also equivalent to the typical δ^{34} S value of (+0.3+/-0.5)‰ found in submarine glasses (Sakai et al., 1984). When uncertainty of δ^{34} S value in vesicles is similar to that obtained by bulk analyses, a δ^{34} S value of (0+/-0.5)‰ can be applied for the vesicle value.

3.3.2. Glass matrix

³He concentrations in MORB glass matrix are listed in Table 2-6. ³He concentrations in RY380-R03b and 418R002 were taken from my own earlier study (Kagoshima et al., 2012). ³He concentrations in glass matrix of other samples were estimated multiplying those in vesicles and glass/vesicle content ratios of ³He in RY380-R03b and 418R002 which had been calculated to be 1.43 as an average of 2.48 and 0.37 for RY380-R03b and 418R002, respectively. The ³He concentrations in glass matrix have the range of (9-57)x10⁻¹⁶ mol/g. The sulfur contents of MORB glass matrix measured with a NanoSIMS listed in Table 2-6 are well within the variation of sulfur contents in MOR glass obtained using a conventional method (Sakai et al., 1984). The average value of S/³He ratios in the glass matrix was $(1.8+/-0.5)x10^{10}$ (1 σ). Observed δ^{34} S values are consistent with those of

Sample name	Number of	³ He/ ⁴ He	³ He concentration	S concentration	S/ ³ He
	analyses	(R _a)	$(10^{-16} mol/g)$	(10 ⁻⁹ mol/g)	(10^{6})
(East Pacific Rise)					
RY380-R03b	4	9.49 ± 0.16	23 ± 1	41 ± 6	18 ± 3
ND 18-1	2	9.57 ± 0.28	6 ± 1	33 ± 2	58 ± 8
CL DR01	3	$8.02 \hspace{0.2cm} \pm \hspace{0.2cm} 0.17$	37 ± 6	88 ± 19	27 ± 11
(Mid-Atlantic Ridge)					
418R002	4	8.29 ± 0.13	51 ± 4	173 ± 21	34 ± 5
CH98 DR12	2	9.22 ± 0.24	10 ± 1	48 ± 0.4	48 ± 4
CH31 DR12	2	$8.82 \ \pm 0.70$	13 ± 6	146 ± 42	$129 \ \pm \ 30$
(Central Indian Ridge)					
KH10-6 DR15-glass	7	$8.13 \pm 0.07 *$	37 ± 5 *	140 ± 14	36 ± 3 **
KH10-6 DR16-glass	4	$7.96 \ \pm 0.09$	28 ± 1	55 ± 4	20 ± 2
Average		8.69 ± 0.24	25 ± 5	90 ± 19	46 ± 13

Table 2-5. Helium isotopic compositions, sulfur concentrations, and S/³He ratios in MORB glass vesicles.

Uncertainty: 1o

Blank contributions have been already subtracted from all values listed in this table.

Numbers of datasets used for calculations of values with superscripts * and ** are five and six, respectively. They are different from "7" listed in the "Number of analyses" column. These data reductions were caused by mistakes during helium isotope analyses.

Sample name	³ He concentration	Number of SIMS	S concentration	S/ ³ He	Number of	$\delta^{34}S$
	(10 ⁻¹⁶ mol/g)	analyses	(10 ⁻⁶ mol/g)	(10 ⁸)	IRMS analyses	(‰)
(East Pacific Rise)						
RY380-R03b	57 *	18	43	76	3	0.45
ND 18-1	9 **	10	42	472		
CL DR01	53 **	15	39	73	1	1.26
(Mid-Atlantic Ridge)						
418R002	19 *	18	34	181	3	0.53
CH98 DR12	14 **	15	41	290		
CH31 DR12	19 **	15	33	176	1	0.78
(Central Indian Ridge)						
KH10-6 DR15-glass	53 **	10	26	49	1	-0.19
KH10-6 DR16-glass	40 **	10	38	94	1	0.11
Average	33		37	176		0.49

Table 2-6. ³He and sulfur concentrations, S/³He ratios, and δ^{34} S values of MORB glass matrix.

*: ³He concentrations reported by Kagoshima et al. (2012).

**: ³He concentrations estimated based on those in vesicles and glass/vesicle content ratios of ³He in RY380-R03b and 418R002. See text for the calculations.
MOR basalt and mantle values (Sakai et al., 1984; Marini et al., 2011), suggesting the typical upper mantle-derived sulfur signature.

3.4. Discussion

3.4.1. Representative components derived from the upper mantle at mid-ocean ridges

The representative S^{3} He ratio and the ³He flux derived from the upper mantle at MOR are needed for sulfur flux calculations. The observed S/³He ratios of vesicles in MORB glass range from 1.8×10^7 to 1.3×10^8 (Table 2-5). These values are lower than 4.9×10^9 - 4.7×10^{10} found in the glass matrix (Table 2-6). This observation suggests higher solubility of sulfur than of helium in basaltic melt, which is also supported by recent laboratory experiments (Webster & Botcharnikov, 2011). The average S/³He ratio in MORB glass vesicles is $(4.6+/-1.3)\times 10^7$. Although vesicle components are preferentially released from magmas by propagations of film-rupturing as discussed above, some parts of sulfur in voids are thought to be finally trapped as sulfide spherules on vesicle walls (Moore & Calk, 1991). Such sulfide spherules are not dissolved into alkaline solutions those I used for experiments, thus the average vesicle S^{3} He ratio may be considered as the minimum value derived from the upper mantle at MOR. Another possible way to determine a representative $S/^{3}$ He ratio at MOR is to study chemistry of submarine hydrothermal fluids. The average $S^{/3}$ He ratio is $(3.4+/-0.7)x10^8$ (1 σ) among 10 high-temperature (>200C) hydrothermal sites worldwide (Table 2-7). The δ^{34} S value of H₂S in hot vent fluids is variable (Marini et al., 2011), but the original value before the incursion of seawater is similar to the MOR basalt values (Alt, 1995; Gamo, 1995). A small part of H₂S might be generated from the reduction of seawater SO₄ from the recharge zone (Alt, 1995), although it is difficult to deconvolve the contribution from the mantle and seawater precisely. Therefore the vent $S^{/3}$ He ratio of $(3.4+/-0.7)x10^8$ is expected to be the maximum estimate at MOR. One might take an average of these two independent estimates $(1.9+/-1.5) \times 10^8$ as the representative value derived from the upper mantle at MOR. S^{3} He ratios in MORB glass matrix are higher than this upper mantle value (Table 2-6), implying that helium has degassed from the melt before it was quenched to glasses and therefore these ratios are not appropriate to be considered for sulfur flux calculations.

Table 2-7. Helium isoto	opic composit	ions, H_2S c	concentrati	ons, and ĉ	3 ³⁴ S value	s of H ₂ S i	n high-temperature (>200C) hydrothermal fluids from mid-ocean ridges.
Hydrothermal fluids	Temperature (C)	³ He/ ⁴ He (R _a)	³ He (pM)	H ₂ S (MM)	δ ³⁴ S (%o)	S^{3} He (10 ⁷)	Reference
(Eastern Pacific)					r.		
Juan de Fuca (47-48°N)	370	8.0	17.4	4.4		25.3	Butterfield et al. (1994)
East Pacific Rise (21°N)	340	7.8	17.3	7.5	3.1	43.4	Woodruff & Shanks (1988); Welhan & Craig (1983); Kim et al. (1984); Campbell et al. (1988a)
East Pacific Rise (13°N)	380	7.5	24.7	5.2		21.0	Kim et al. (1984); Merlivat et al. (1987); Bowers et al. (1988)
East Pacific Rise (11°N)	347	8.3	18.2	8.2	4.7	45.1	Kim et al. (1984); Bowers et al. (1988); Bluth & Ohmoto (1988)
East Pacific Rise (17-19°S)	382	8.3	9.2	7.0	6.2	75.4	Shanks (2001); Jean-Baptiste et al. (1997)
(Mid-Atlantic Ridge)							
Menez Gwen (38°N)	284	8.7	17.6	1.6		9.1	Charlou et al. (2000)
Lucky Strike (37°N)	324	8.1	4.5	2.7		61.0	Charlou et al. (2000)
Rainbow (36°N)	365	7.5	25.0	1.2		4.8	Charlou et al. (2002); Jean-Baptiste et al. (2004)
TAG (26°N)	358	7.5	15.8	4.0	8.5	25.4	Charlou et al. (1996); Rudnicki & Elderfield (1992); Chiba et al. (2001)
Snake Pit (23°N) MARK	350	7.9	19.4	5.9	5.0	30.5	Rudnicki & Elderfield (1992); Campbell et al. (1988b)
Average		8.0	16.9	4.8	5.5	34.1	

3.4.2. Sulfur flux at mid-ocean ridges

As described earlier, based on the saturation anomaly of ³He in deep seawater of the eastern Pacific, a value of (1070+/-270) mol/y was calculated for the ³He flux at MOR (Craig et al., 1975), but a more recent estimate of the MOR ³He flux is (527+/-102) mol/y derived from an ocean circulation model which also considers radiocarbons and chlorofluorocarbons together with ³He anomaly in deep seawater (Bianchi et al., 2010). The combination of this MOR ³He flux of (527+/-102) mol/y and the average S/³He ratio of $(1.9+/-1.5)\times 10^8$ obtained in this study provides MOR sulfur flux of $(1.0+/-0.8)\times 10^{11}$ mol/y (Fig. 2-5). This flux is depicted in Fig. 2-6 together with previously reported fluxes. This mantle flux is consistent with the estimate based on seawater-basalt sulfur isotopic exchange during hydrothermal alteration (Alt, 1995), but it is about an order of magnitude smaller than the value calculated by the production rate of the oceanic crust and sulfur contents therein (Hansen & Wallmann, 2003). This difference suggests that most sulfur remains in the silicate melt and finally solidifies as sulfide in the oceanic crust, and does not contribute to the MOR flux that dissolves and discharges into the ocean. This is consistent with suggestions by Sakai & Matsuhisa (1996) and Tajika (1998), indicating a small part of sulfur in the oceanic crust is released into the ocean, while the most part remains in the oceanic crust. When we consider the mass balance of sulfur in global ocean water, the mantle sulfur flux of $(1.0+/-0.8)\times 10^{11}$ mol/y is a second-order flux compared with the riverine input of 8.9×10^{11} mol/y into the ocean and the output of 5.5×10^{11} mol/y as sedimentary pyrite and evaporitic sulfate (Holser et al., 1988). However, the sulfur flux from the upper mantle into the ocean represents a deep Earth contribution and is therefore distinct from riverine sulfur input that is continent-derived and may be strongly related to return from subducting materials.

3.4.3. Sulfur flux at arc volcanoes

ARC ³He flux was estimated from the MOR flux, given the assumption that the magma production rate of ARC is 20% of that of MOR. This portion is consistent with the estimate of global magma emplacement and volcanic output averaged over the last 180 m.y. (Crisp, 1984). Currently accepted MOR ³He flux of (527+/-102) mol/y (Bianchi et al., 2010) would result in an ARC ³He flux of (105+/-20) mol/y which is consistent with the value obtained by summation of ³He flux at arc volcanoes worldwide (Hilton et



Figure 2-5. Global sulfur fluxes among the Earth's interior and surface reservoir.

Numbers are sulfur fluxes in the unit of Gmol/y among reservoirs with errors as subscripts. See text for calculations of fluxes at mid-ocean ridges (100 Gmol/y) and arc volcanoes (680 Gmol/y), and the mantle contribution to the ARC flux (21 Gmol/y). Other fluxes related to recycle and return processes were calculated under the assumption of the steady-state surface reservoir of sulfur.





Sulfur fluxes at MOR are from this study, Hilton et al. (2002), Hansen & Wallmann (2003), Sakai & Matsuhisa (1996), Tajika (1998), and Alt (1995). The residence times were calculated using sulfur fluxes and sulfur surface inventory of 5.3x10²⁰ mol/y reported by Hilton et al. (2002). The Earth's age of 4.54x10⁹ years (Tera, 1981) is depicted as the dashed line. The residence time longer than the Earth's age obtained in this study allows a catastrophic degassing history of sulfur. See text for the errors.

Momo	1 24	Ion	Location	Temp.	³ He/ ⁴ He	$\delta^{34}S$	$S/^{3}He$	Deferences
INALLIC	Lál.	L011.	госацон	(°C)	(R_a)	(%0)	(10^{9})	Neterence
Koryak	53.32N	158.69E	Kamchatka	215	5.3		8.0	Taran (2009)
Avacha	53.26N	158.83E	Kamchatka	473	7.1	6.4	2.3	Taran (2009); Taran et al. (1997)
Mutnovsky	52.45N	158.2E	Kamchatka	543	8.2	7.7	10.8	Taran (2009); Taran et al. (1992)
Kudryavy	46N	148E	Kuril	912	6.8	4.7	1.6	Taran (2009); Taran et al. (1995)
Usu	42.53N	140.83E	Japan	750	5.4	7.0	3.4	Fischer (2008); Marty et al. (1989); Kasasaku et al. (1999)
Kuju	33.1N	131.3E	Japan	580	7.6	4.2	7.8	Mizutani et al. (1986); Nagao et al. (1981); Saito et al. (2002)
Unzen	32.75N	130.3E	Japan	818	7.3		2.1	Fischer (2008); Sano & Williams (1996)
Satsuma-Iwojima	30.8N	130.3E	Japan	885	7.9	11.7	12.6	Taran (2009); Kasasaku et al. (1999)
Merapi	7.52S	110.43E	Indonesia	803	7.4		3.1	Taran (2009)
Lewotolo	8.27S	123.5E	Indonesia	490	3.6	5.2	11.6	Poorter et al. (1991); Varekamp et al. (1992)
White Island	37.52S	177.18E	New Zealand	495	6.1	4.0	8.4	Marty & Giggenbach (1990); Giggenbach (1982)
Ngauruhoe	39.15S	175.63E	New Zealand	640	6.7	4.4	8.4	Fischer (2008); Giggenbach (1982); Giggenbach et al. (1993); Giggenbach (1996)
Momotombo	12.41N	86.52E	Nicaragua	747	7.1	5.9	4.6	Taran (2009); Menyailov et al. (1986)
Galeras	1.22N	77.37W	Colombia	642	8.4	-8.9	1.0	Marini et al. (2011); Fischer (2008); Sano & Williams (1996)
Colima	0.95N	77.89W	Mexico	814	6.6	3.4	12.2	Taran (2009); Taran et al. (2001)
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al., 2002). I selected 15 high-temperature (>200C) volcanic gases around the world and compiled their helium and sulfur compositions in Table 2-8. Their ${}^{3}\text{He}/{}^{4}\text{He}$ ratios range from 3.6 R_a to 8.4 R_a which are consistent with the range of subduction-type helium signature (Hilton et al., 2002; Sano & Fischer, 2013). S/ ${}^{3}\text{He}$ ratios in volcanic gases range from 1.0x10⁹ to 1.3x10¹⁰ with an average value of (6.5+/-1.1)x10⁹ (1 σ). Therefore, the ARC sulfur flux is estimated to be (6.9+/-1.7)x10¹¹ mol/y multiplying this volcanic gas ratio and the ARC ${}^{3}\text{He}$ flux of 105+/-20 mol/y.

3.4.4. Deconvolution of sulfur flux at ARC

In order to estimate the sulfur flux from the mantle to the surface through arc volcanisms, I deconvolved magmatic sulfur into the mantle component and other endmembers. As discussed for carbon, sulfur and helium isotopic compositions are also useful for investigating the origin of sulfur at arc volcanoes. The thermodynamic equilibrium between SO₂ and H₂S together with their δ^{34} S values might provide constraints on the evolution of volcanic gases, such as an isochemical cooling path, under the assumption that the initial $\delta^{34}S_{\Sigma S}$ value is (0+/-0.5)‰ as discussed for the results of vesicle analyses, where $\delta^{34}S_{\Sigma S}$ denotes the total sulfur isotopic ratio of SO₂ and H₂S. In addition, the $\delta^{34}S_{\Sigma S}$ values might provide information related to the origin of sulfur in ARC volcanic gases, even though they might be affected by a gas-melt separation and related fractionation processes (Marini et al., 2011; de Moor et al., 2013). Available data of ³He and total sulfur contents, and $\delta^{34}S_{\Sigma S}$ values for high-temperature volcanic gases (>200C) in subduction zones were compiled from the literature and listed in Table 2-8. The $\delta^{34}S_{\Sigma S}$ values are generally positive, except for one outlier from Galeras. This similarity suggests that the sulfur signature of an ARC magma source is due to incorporation of subducted sulfate partly derived from a seawater component (de Moor et al., 2013; Sakai & Matsubaya, 1977; Alt et al., 1993) with high δ^{34} S values. The average S^{3} He ratio of $(6.5+/-1.1)\times 10^{9}$ among these high-temperature ARC gases is significantly high value relative to that of the upper mantle, suggesting enrichment of sulfur in the ARC mantle source by subduction processes.

As discussed for carbon, following Sano & Marty (1995), sulfur and helium isotopic compositions in arc volcanic gases can be explained by some-component mixing because





Volcanic gas data described in Table 2-8 and 2-9 are shown as red circles. Three square zones are end-members of volcanic gases: the upper mantle-derived component, sedimentary pyrite with reduced sulfur derived from slab abbreviated as "sedimentary pyrite", and subducted sulfate. Mixing lines of three end-members are depicted as solid curves.

Name	$\delta^{34}S$	S / ³ Ha	Sulfur source co	ontribution	
Name	(‰)	5/ He	Upper mantle	Sedimentary pyrite	Subducted sulphate
Koryak		8.0E+09	2.4%		
Avacha	6.4	2.3E+09	8.4%	29.9%	61.7%
Mutnovsky	7.7	1.1E+10	1.8%	29.3%	68.9%
Kudryavy	4.7	1.6E+09	11.9%	33.1%	55.0%
Usu	7.0	3.4E+09	5.7%	29.3%	65.0%
Kuju	4.2	7.8E+09	2.4%	39.7%	57.9%
Unzen		2.1E+09	9.1%		
Satsuma-Iwojima	11.7	1.3E+10	1.5%	17.1%	81.4%
Merapi		3.1E+09	6.1%		
Lewotolo	5.2	1.2E+10	1.6%	37.1%	61.3%
White Island	4.0	8.4E+09	2.2%	40.5%	57.3%
Ngauruhoe	4.4	8.4E+09	2.3%	39.3%	58.4%
Momotombo	5.9	4.6E+09	4.2%	33.5%	62.4%
Galeras	-8.9	1.0E+09	18.8%	71.4%	9.8%
Colima	3.4	1.2E+10	1.6%	42.6%	55.9%
Average values	4.6	6.5E+09	2.9%	38.1%	59.0%
(End-members)					
Upper mantle	0.0	1.9E+08	100.0%		
Sedimentary pyrite	-14.9	1.0E+13		100.0%	
Subducted sulfate	17.5	1.0E+13			100.0%

Table 2-9. Sample and end-member compositions with sulfur contributions.

Data of subaerial volcanoes and three end-members are described here. See text for the calculations of end-member contributions.

sources of volcanic gases are mixtures of the upper mantle component and subducting materials. I apply here the upper mantle-derived component, sedimentary pyrite with reduced sulfur derived from slab (abbreviated as sedimentary pyrite hereafter), and subducted sulfate as end-members producing volcanic gas compositions (Fig. 2-7; Table 2-9). The S/³He ratio derived from the upper mantle is $(1.9+/-1.5)\times 10^8$ determined by compositions of MORB vesicles and submarine hydrothermal fluids in the current study. As carbon, one would apply a fixed S^{3} He ratio of $10^{(13+/-1)}$ for subducting materials such as sedimentary pyrite, slab components, and subducted sulfate because pyrite and slab are depleted in the primordial ³He injected from the mantle (Hiyagon, 1994), and the seawater $SO_4/^3$ He of $1.0x10^{14}$ is almost equivalent to the value suggested here for subducting materials. The δ^{34} S values of sedimentary pyrite vary considerably as a result of bacterial reduction of seawater sulfate, and have a mean value of -20.9‰ in the Western Pacific (Alt & Burdett, 1992). Results of a recent study (Alt et al., 2012) of the oceanic basement in northern Italy suggest that low-temperature serpentinization produces a negative $\delta^{34}S_{\Sigma S}$ value with (-8.9+/-8.0)‰. An average of the two components, (-14.9+/-6.0)‰, can be regarded as a representative δ^{34} S value of sedimentary pyrite. On the other hand, subducted sulfate is mostly derived from sulfate dissolved in seawater with a δ^{34} S value of +21.0‰ (Rees, 1978). Metasomatic fluids released from sediment, of which the sulfur is mostly in the form of sulfate, have a δ^{34} S value of +14‰ when their sulfur compositions resemble the bulk sediment composition (Alt et al., 1993). Using these values, the δ^{34} S value of subducted sulfate is defined as (+17.5+/-3.5)% as an average of the two components. Figure 2-7 presents the relation between S^{3} He ratios and $\delta^{34}S_{\Sigma S}$ values in arc volcanic gases, indicating that their sulfur and helium compositions can be exactly explained by mixing of three components applied here.

When sulfur in volcanic gases are mixtures of the upper mantle, sedimentary pyrite, and subducted sulfate having respective masses M, P, and S, contribution percentages of the three components in the ARC samples can be deconvolved solving the following equations:

$$\delta^{34}S_{V} = \delta^{34}S_{M}M + \delta^{34}S_{P}P + \delta^{34}S_{S}S$$
(2-4)

$$1/(S^{3}He)_{V} = M/(S^{3}He)_{M} + P/(S^{3}He)_{P} + S/(S^{3}He)_{S}$$
 (2-5)

 $M + P + S = 1 \tag{2-6}$

In those equations, subscripts V, M, P and S denote the volcanic gases (ARC samples), the upper mantle, sedimentary pyrite, and subducted sulfate, respectively. As discussed above, $\delta^{34}S_M = (0+/-0.5)\%$, $\delta^{34}S_P = (-14.9+/-6.0)\%$, $\delta^{34}S_S = (+17.5+/-3.5)\%$, $(S/^3He)_M = (1.9+/-1.5)x10^8$, $(S/^3He)_P = 1.0x10^{(13+/-1)}$, and $(S/^3He)_S = 1.0x10^{(13+/-1)}$ are applied for contribution calculations. The contribution of mantle sulfur is 1.5%-19% (2.9% average) in ARC samples, and the main contribution derives from subducted sulfate and sedimentary pyrite (Table 2-9). The volcanic gas from Satsuma-Iwojima shows the highest subducted sulfate contribution of 81.4% with the highest $\delta^{34}S_{\Sigma S}$ value of 11.7%. In order to explain the heavy $\delta^{34}S_{\Sigma S}$ values of ARC volcanic gases, incorporation of a seawater component in the magma source has been inferred since the 1970s (Sakai & Matsubaya, 1977). Here we first provide a quantification of the relative amount of the subducted seawater sulfate contribution to ARC gases, which allows us to evaluate the recycling capacity of ARC volcanoes within the global sulfur cycle.

The ARC sulfur flux was calculated to be $(6.9+/-1.7)\times10^{11}$ mol/y as discussed in the previous section. This value is significantly higher than MOR sulfur flux of $(1.0+/-0.8)\times10^{11}$ mol/y also estimated in this study. However the upper mantle contribution to ARC volcanic gases is only (2.9+3.3/-2.4)% of total sulfur, on average. The uncertainty of the upper mantle contribution was calculated based on the highest and the lowest upper mantle contributions calculated from extremely high or low values' sample and end-member compositions as follows: the upper mantle contribution of 6.2% derived from the case applying S/³He ratios of 5.4×10^9 , 3.4×10^8 , 1.0×10^{14} , and 1.0×10^{14} for ARC samples, the upper mantle, sedimentary pyrite, and subducted sulfate; and the upper mantle contribution of 0.5% derived from the case applying S/³He ratios of 7.6×10^9 , 4.0×10^7 , 1.0×10^{12} , and 1.0×10^{12} for ARC samples, the upper mantle, sedimentary pyrite, and subducted sulfate. The sulfur flux from the wedge mantle at ARC then becomes (2.0+2.3/-1.7) $\times10^{10}$ mol/y, which is less than the mantle sulfur flux discharging into the ocean at MOR. The major contribution of the ARC sulfur flux is derived from subducted sedimentary pyrite and subducted sulfate partly derived from the seawater component.

3.4.5. Contribution to global flux of sulfur from present hot spot magmatism

Intra-plate volcanism such as that at Hawaii or Yellowstone might release large amounts of sulfur, but no study for that has been reported in the literature. The global hot spot ³He flux of 2.5 mol/y was reported based on Hawaiian magma (Torgersen, 1989), which was not scaled to MOR flux as for ARC. It is not necessary to revise the value according to the new MOR flux (Bianchi et al., 2010). The S/³He ratio of Kilauea volcanic gas is 5.6×10^8 , where the S/He ratio and ³He/⁴He ratio were referred from the literature (Craig & Lupton, 1976; Fischer, 2008). Yellowstone and Iceland fluid samples (Hearn et al., 1990; Sano et al., 1985) show ratios of 5.0×10^7 and 1.1×10^8 , which are both smaller than the Hawaiian value, although their temperatures are lower than 200C. Then the global sulfur flux at hot spot regions might not exceed the value of 1.4×10^9 mol/y, which is almost two orders of magnitude smaller than the MOR flux of 1.0×10^{11} mol/y. Therefore the hot spot magmatism does not strongly affect the global sulfur cycle and is ignored for discussion of geochemical sulfur cycles in this study.

3.4.6. Global sulfur cycle

A summary of the global sulfur flux is depicted in Fig. 2-5. Present hot spot magmatism does not likely contribute substantially to the global flux of sulfur as discussed in the previous section. The total volcanic flux of sulfur is estimated as (7.9+/-1.9)x10¹¹ mol/y and represents about one-third of the anthropogenic emissions due to coal burning and sulfide ore smelting (Moller, 1984). This natural flux, if it has remained constant over 4.54 billion years of geological time (Tera, 1981), engenders an accumulation of $(3.6+/-0.9)\times 10^{21}$ mol. This value is greater than the surface inventory of 5.3×10^{20} mol (Hilton et al., 2002). If we take the MOR flux together only with the mantle wedge flux of $(2.0+2.3/-1.7)\times 10^{10}$ mol/y, then the accumulation becomes $(5.4+/-3.8)\times 10^{20}$ mol in total, which is comparable to the surface inventory. When steady-state recycling of sulfur is applied, the total subducting flux becomes $(7.9+/-1.9)\times 10^{11}$ mol/y. This value is lower than 4.11×10^{12} mol/y reported by Hilton et al. (2002). Such contradiction may be derived by the difference in estimation methods of sulfur fluxes, that is, Hilton's flux includes components staying in the oceanic crust and subducted which is not considered as fluxes in this study because such components does not contribute to the global flux from the mantle to the surface.

4. Conclusions

Sulfur and helium compositions in MORB vesicles were determined by the currently established crushing technique based on the conventional noble gas analyses. The advantage of this technique is that sulfur fluxes could be calibrated against the welldetermined ³He fluxes because sulfur is extracted simultaneously with helium isotopes. The data obtained by the crushing method was applied to constrain the sulfur flux at MOR. In conclusion, the best estimates of MOR sulfur and carbon flux are $(1.0+/-0.8)x10^{11}$ mol/y and $(1.2+/-0.4)\times10^{12}$ mol/y, respectively at present, which are less than their volcanic fluxes at ARC. Sulfur and carbon fluxes from only the mantle wedge to the surface environment at ARC were calculated to be (2.0+2.3/-1.7)x10¹⁰ mol/y and (2.3+1.5/-1.2)x10¹¹ mol/y, respectively. Summations of MOR fluxes and mantle contributions to ARC fluxes provided mantle-derived sulfur and carbon fluxes of (1.2+/- $(0.8)\times10^{11}$ mol/y and $(1.4+/-0.4)\times10^{12}$ mol/y, respectively. These data provide a C/S flux ratio of (12+/-8) which is similar to the C/S ratio in the surface inventory of 13 (Hilton et al., 2002). The results suggest that the main source of sulfur and carbon is the upper mantle. To balance the mass between the crust and the mantle, the sulfur subducted into the mantle and not immediately recycled to the surface is expected to be equivalent to $(1.2+/-0.8)\times10^{11}$ mol/y, which is about (15+/-11)% of the subducted sulfur of (7.9+/-1.9)x10¹¹ mol/y. We calculated sulfur and carbon fluxes from the mantle based on the plausible S^{3} He and C^{3} He ratios and the recently reported ³He flux at MOR, which constrained geochemical cycles of sulfur and carbon, and evolutionary histories of the atmosphere and hydrosphere.

Chapter 3 Geochemical cycles of halogens

Contents in this chapter will be published in five years.

本章については、5年以内に雑誌等で刊行予定のため非公開。

Chapter 4 Summary and general conclusions

Earth's atmosphere and ocean have been produced by accumulation of volatile species injected from the mantle through volcanic and hydrothermal activity. Sulfur and carbon are major components of volcanic gases which are injected into surface environment in forms of SO₂, H₂S, CO₂, CH₄, and they significantly affect atmospheric and oceanic chemistry and their evolution. Therefore their fluxes are important to reveal Earth's evolution and to assess how much volcanic activity affects chemistry of surface environment. Halogens are also important volatile elements for elucidating evolutionary histories of Earth's atmosphere and ocean because they form variable chemical compounds in surface environments due to their high chemical activity. In addition, they are concentrated in fluids and have been used as hydrogeological tracers. In order to constrain their evolution on Earth, I have assessed their behavior between the mantle and surface environments in this study. For sulfur and halogen fluxes, an approach using calibration against the reliable ³He flux was applied for elucidating their cycles.

In *Chapter 2*, in order to estimate sulfur flux at mid-ocean ridges (MOR), I focused on the volatile components preferentially released from magmas through film rupturing processes in ascending magmas. Such components are expected to be retained in vesicles of MORB glasses, and they are analyzed with the currently established crushing technique based on the conventional noble gas analyses. Sulfur is extracted into the solution and measured using ion chromatography, while helium isotopes are measured using a noble gas mass spectrometer in the protocol. The advantage of this technique is that sulfur fluxes could be calibrated against the well-determined ³He fluxes because sulfur is extracted simultaneously with helium isotopes. The data obtained by the crushing method was applied to constrain the sulfur flux at MOR. Carbon flux was also calibrated against the ³He flux. In conclusion, the best estimates of MOR sulfur and carbon fluxes are $(1.0+/-0.8)x10^{11}$ mol/y and $(1.2+/-0.4)x10^{12}$ mol/y, respectively at present, which are less than their volcanic fluxes at arc volcanoes (ARC). Wedge mantle contributions to ARC fluxes have been deconvolved, providing total mantle fluxes combined with MOR results. I calculated the mantle C/S flux ratio as (12+/-8) which is similar to the C/S ratio in the surface inventory of 13. The results suggest that the main source of sulfur and carbon is the upper mantle.

In *Chapter 3*, in order to estimate halogen fluxes at MOR, I focused on the volatile components preferentially released from magmas as sulfur, and analyzed MORB glasses. The complete halogens and helium datasets in MORB vesicles were obtained by combinations of some independent analyses. Helium, fluorine, and chlorine data were obtained by crushing experiment followed by noble gas mass spectrometry and ion chromatography, which is the similar way as sulfur. Chlorine, bromine, and iodine contents in vesicles were determined by the noble gas method which converts halogens into proxy noble gas isotopes. Combinations of data enabled me to estimate halogen fluxes at MOR by calibrating against the well documented ³He flux.

This part will be published in five years.

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In conclusions, I have determined sulfur, carbon, and halogen fluxes between the mantle and surface environment by calibrating against the reliable ³He fluxes. Simultaneous analyses of helium and other highly reactive volatile elements, when combined with other conventional analytical techniques, provided complete datasets of sulfur, halogens and helium.

This part will be published in five years. この部分については、5年以内に雑誌等で刊行予定のため非公開。 I have investigated geochemical cycles of variable volatile elements in my Ph.D. course followed by a half year as a research associate in Atmosphere and Ocean Research Institute. In the future, research focusing on behaviors of volatile elements including helium at individual volcano is needed in order to elucidate volatile cycles in subduction zones more precisely. This is also important to assess volcanic hazards such as enormous eruptions. A case study has been conducted at Ontake volcano for revealing helium and carbon behaviors related to eruptions (see *Appendix C*), indicating that helium isotopes are the most sensitive magmatic tracers and useful to assess volcanic activity. Chlorine isotopes could be also useful tracers for investigating volatile cycles at subduction zones. Chlorine isotope analyses with reasonably high resolution using NanoSIMS have been conducted (see *Appendix B*), and the results showed that the NanoSIMS technique is useful for assessing aqueous alteration. The precision should be enhanced to trace magmatic chlorine for elucidating geochemical cycles of halogens and water.

Appendix A Frozen Crushing Method

Sulfur and halogen concentrations in MORB glasses were measured for estimating their mantle fluxes and geochemical cycles in this study. The detail of the crushing method (Frozen Crushing Method, abbreviated as FCM hereafter; Kagoshima et al., 2012) applied to extract sulfur, fluorine, chlorine, and noble gases from MORB vesicles simultaneously in *Chapter 2* and *3* is described below. Applying FCM, MORB vesicle compositions are determined by crushing with frozen alkaline solutions followed by analyses using a noble gas mass spectrometer and an anion chromatography system.

FCM is a modified version of the conventional in vacuo crushing method for solid samples. A difference is that samples are loaded in the vacuum crushing vessel with alkaline solutions such as NaOH solution and NH₃ solution. 0.5g to 1.0g of fresh submarine glass aliquots were selected by eyes and washed by ultra-sonic bath with Milli-Q water (>18 M Ω cm; Millipore Co.) for five minutes followed by ethanol for five minutes. After that, ultra-sonicated samples were loaded into a stainless-steel crusher with 1-2 cm³ of alkaline solution (diluted NaOH or NH₃ solutions whose concentrations were 0.2-4.0 mol/L and 5.6 mol/L (10%), respectively) and a stainless-steel ball for crushing (Fig. A-1b). In the conventional way, air fraction is removed by vacuum before crushing because atmospheric noble gases can easily mask sample noble gas signatures. However, FCM does not allow us to pump out air in the room temperature because the vessel contains liquid water inside. This problem can be solved by pumping out after freezing solutions at liquid nitrogen temperature (77K). Closing the crusher, the alkaline solution was frozen by cooling the crusher itself in liquid nitrogen, and the air in the crusher were removed using a rotary pump (Fig. A-1c). Once the frozen solution was melted at the room temperature using hot water, it was frozen again using liquid nitrogen and evacuated with a turbo molecular pump (TMP) for five minutes. Applying this TMP evacuation procedure to the solution for two times, the atmospheric noble gases dissolved in the solution and on the sample surface were removed, and blanks of helium were resulted to be negligibly small compared with actual samples. The blank contribution to ³He concentrations in samples were calculated as less than 1%. After the TMP procedure, sample aliquots were crushed by 1500 to 2000-time strokes of the vessel at around liquid



Figure A-1. Frozen Crushing Method.

(a) A stainless-steel crusher and a stainless-steel ball used for FCM; (b) a schematic view of inside of the crusher; (c) the crusher attached to a vacuum line being cooled at the temperature of 77K by liquid nitrogen; (d) MORB glass aliquots loaded into the crusher; (e) a SEM image of glass vesicles where volatiles are extracted from; (f) a noble gas mass spectrometer (VG5400) used for helium isotope analyses and the vacuum line for purification of gases transferred from the crusher with vectors indicating flows of purification; and (g) an ion chromatography system (ICS-2100) used for sulfur and halogen analyses of solutions extracted from the crusher.

nitrogen temperature which keeps the solution mostly frozen. The glass aliquots were crushed with the frozen solution by the stainless-steel ball in the crusher, and volatiles were extracted from the glass vesicles by mechanical fracturing. Within the extracted volatiles, sulfur and halogen species, which are highly reactive ones, are fixed immediately in the semi-frozen alkaline solution. On the other hand, noble gases including helium isotopes are extracted into the gaseous phase in the crushing vessel. Adsorption of much amount of the solution on the inner wall of the crushing vessel was avoided by the low temperature keeping the solution mostly frozen.

After the crushing procedures, the solution was frozen again and the gaseous phase in the crusher was introduced into a vacuum line for purification followed by noble gas analyses (Fig. A-1f). At first, H₂O and CO₂ were trapped by a U-tube with the liquid nitrogen temperature (77K) for 1-2 minutes. Subsequently, other major components such as N₂ and O₂ were removed using an activated charcoal also at the liquid nitrogen temperature of 77K for 20 minutes. Next, H₂ and heavy noble gases were removed by a heated Ti-getter and another activated charcoal kept at 77K for 20 minutes. After that, I measured ⁴He/²⁰Ne ratios of the purified gas with a quadrupole mass spectrometer in order to evaluate contamination from air fractions. At last, neon was removed by a cryogenic charcoal kept at 40K. The purified helium gas was introduced into a noble gas mass spectrometer (VG5400; Waters Corp.) at the Atmosphere and Ocean Research Institute (AORI), the University of Tokyo for helium isotope analyses. Using VG5400, the ⁴He amount and the ³He/⁴He ratio in the gas was measured (Sano & Wakita, 1988), and they were calibrated against air and our in-house standard gas which has a particular ³He/⁴He ratio of 20.4 R_a (Helium standard of Japan; Matsuda et al., 2002).

Subsequently, the vacuum in the crusher was broken and the alkaline solution was filtered to remove glass fragments whose diameters are longer than 0.2 μ m. After that, all the sulfur compounds were converted into sulfate ions by oxidation with 1-2 cm³ of hydrogen peroxide (30%). Sulfur, fluorine, and chlorine concentrations in the solution were measured using an ion chromatography system (ICS-2100; Thermo Fisher Scientific Inc.; Fig. A-1g) at AORI. The anions (F⁻, Cl⁻, and SO₄²⁻) were analyzed using an AS11-HC anion exchange column with an AG11-HC guard column (Thermo Fisher Scientific Inc.). An EGCIII-KOH cartridge (Thermo Fisher Scientific Inc.) was used to produce potassium hydroxide eluent. In addition, a CRD-200 carbonate removal devise (Thermo

Fisher Scientific Inc.) was used to remove carbonate ions which interfere sulfate peaks. Blank solutions were prepared by the same concentrations of the alkaline solutions and times of stroking as actual samples. After noble gas, sulfur, and halogen analyses, finely crushed sample glass powders whose diameters are shorter than 212 μ m were weighed in order to determine weights of samples crushed and concentrations of volatile elements originally in the samples. More details about FCM are presented in an earlier report (Kagoshima et al., 2012).

Appendix B Chlorine isotope analyses using NanoSIMS

Contents in this appendix will be published in five years.

本補遺の内容については、5年以内に雑誌等で刊行予定のため非公開。

Appendix C

Helium and carbon isotope variations at Ontake volcano, Japan

Contents in this appendix will be published in five years.

本補遺の内容については、5年以内に雑誌等で刊行予定のため非公開。

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